Equations of motion. For example:

$$(\partial_t + \mathbf{u} \cdot \nabla) \mathbf{u} + 2\Omega \times \mathbf{u} + \frac{\nabla p}{\rho} - \mathbf{g} = \nu \nabla^2 \mathbf{u}$$
$$(\partial_t + \mathbf{u} \cdot \nabla) \theta = \kappa \nabla^2 \theta$$
$$(\partial_t + \mathbf{u} \cdot \nabla) S = \mu \nabla^2 S$$

and bndry / initial conditions. This is really old stuff.

Our job is numeric approx solutions ("modeling")

Equations of motion. For example:

$$(\partial_t + \mathbf{u} \cdot \nabla) \mathbf{u} + 2\Omega \times \mathbf{u} + \frac{\nabla p}{\rho} - \mathbf{g} = \underline{v} \nabla^2 \mathbf{u}$$
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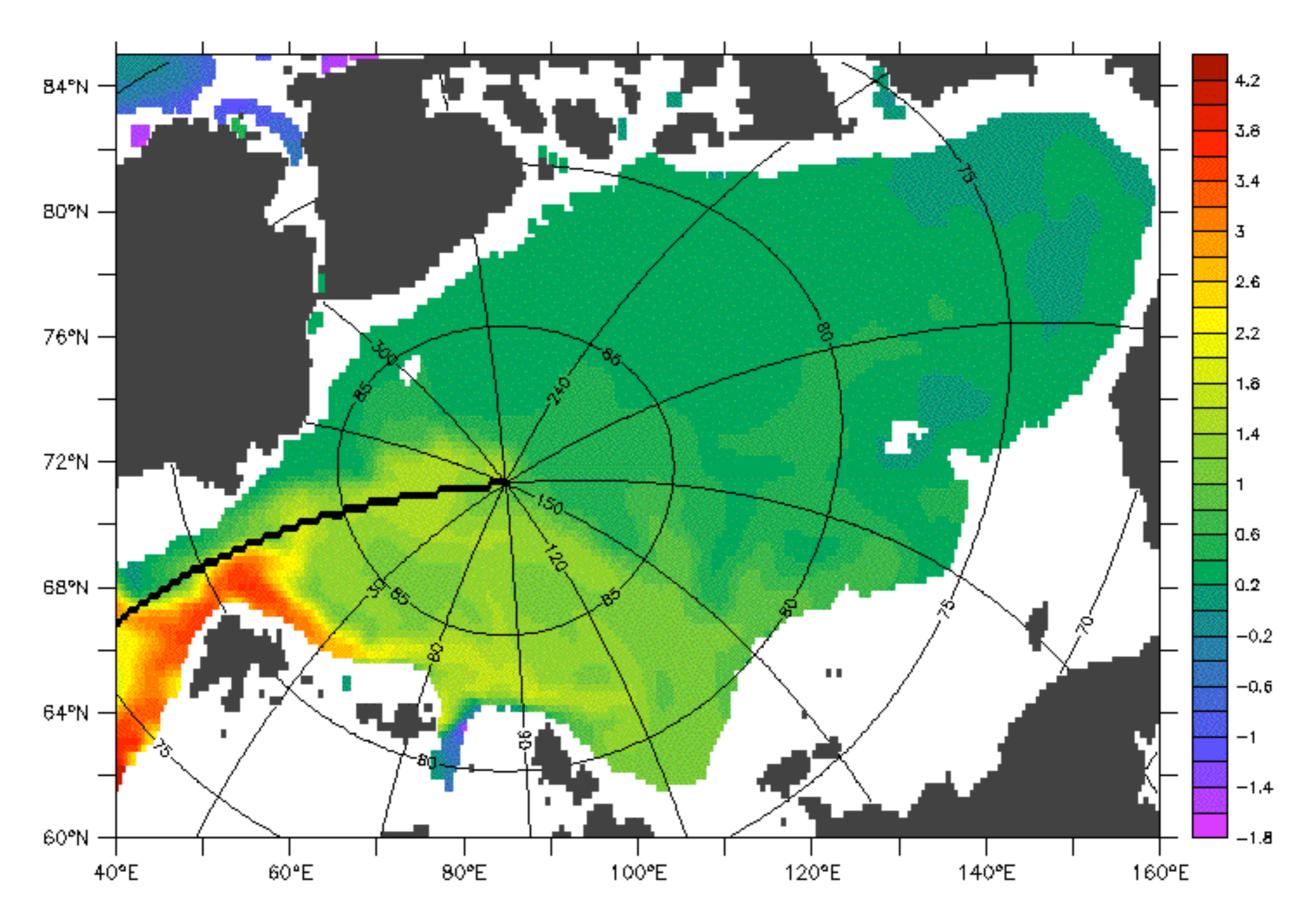
and bndry / initial conditions. This is really old stuff.

Our job is numeric approx solutions ("modeling") -- maybe fudging some coefs ("parameterization").

This talk: Rethink equations of motion. Why? Consider dependent variables (stuff that "moves")?

Hint: it matters.

What is this stuff? θ at some x,y,z,t? How about $\langle \theta \rangle$?



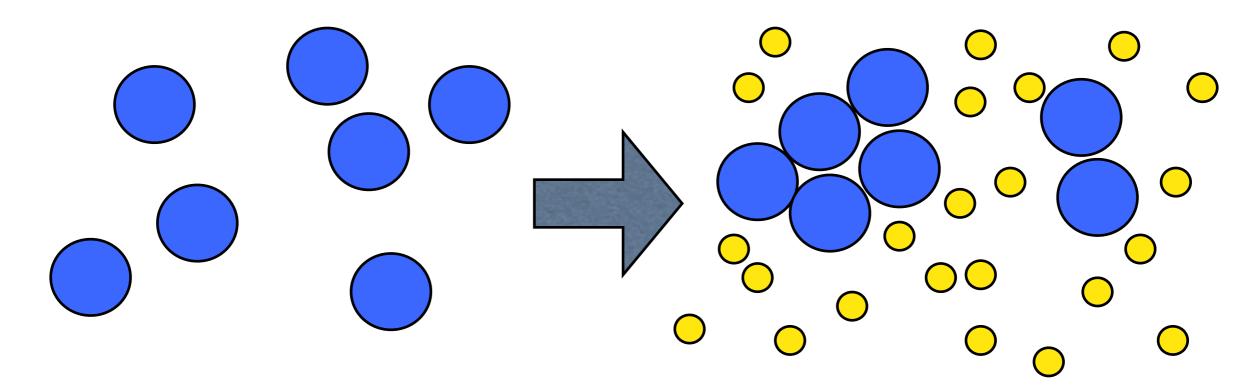
Rewrite symbolically $\frac{\partial \mathbf{Y}}{\partial t} = \mathbf{F}(\mathbf{Y}) + \mathbf{G}$

where Y collects everything that interests you.

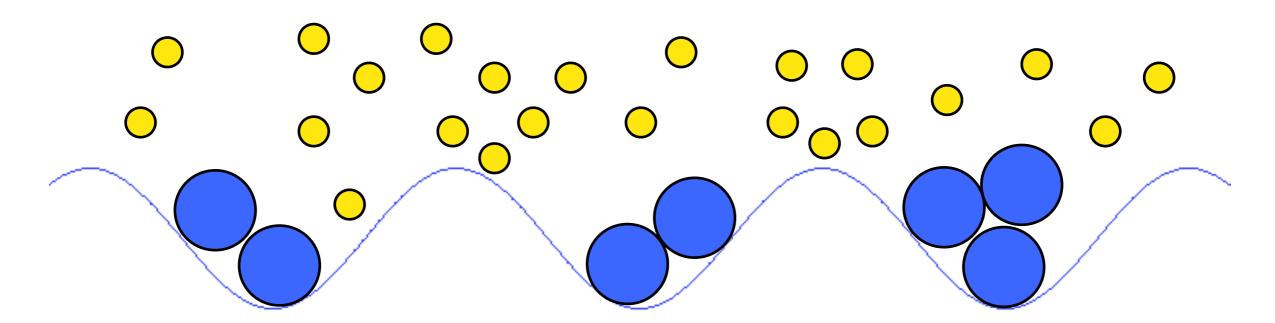
This equation, solved honestly, requires mm grid. Ignore small scale details => large scale expectations.

Probability!
$$dP = P(\mathbf{Y})d\mathbf{Y}$$
 $\langle \mathbf{Y} \rangle \equiv \int \mathbf{Y} dP$
 $\partial \langle \mathbf{Y} \rangle / \partial t = \langle \mathbf{F}(\langle \mathbf{Y} \rangle) \rangle + \langle \mathbf{G} \rangle + "oops"$ What to do
about "oops"?
Entropy: $H \equiv -\int dP \log P$
then $\partial \langle \mathbf{Y} \rangle / \partial t \approx \langle \mathbf{F}(\langle \mathbf{Y} \rangle) \rangle + \langle \mathbf{G} \rangle + \mathbf{C} \cdot \partial_{\langle \mathbf{Y} \rangle} H$

entropic force



Examples from nanoworld (colloids, 'machines', microbiol): The only explicit physics is repulsion among balls, and from walls. "See" attraction. "Entropic forcing" in the lab!

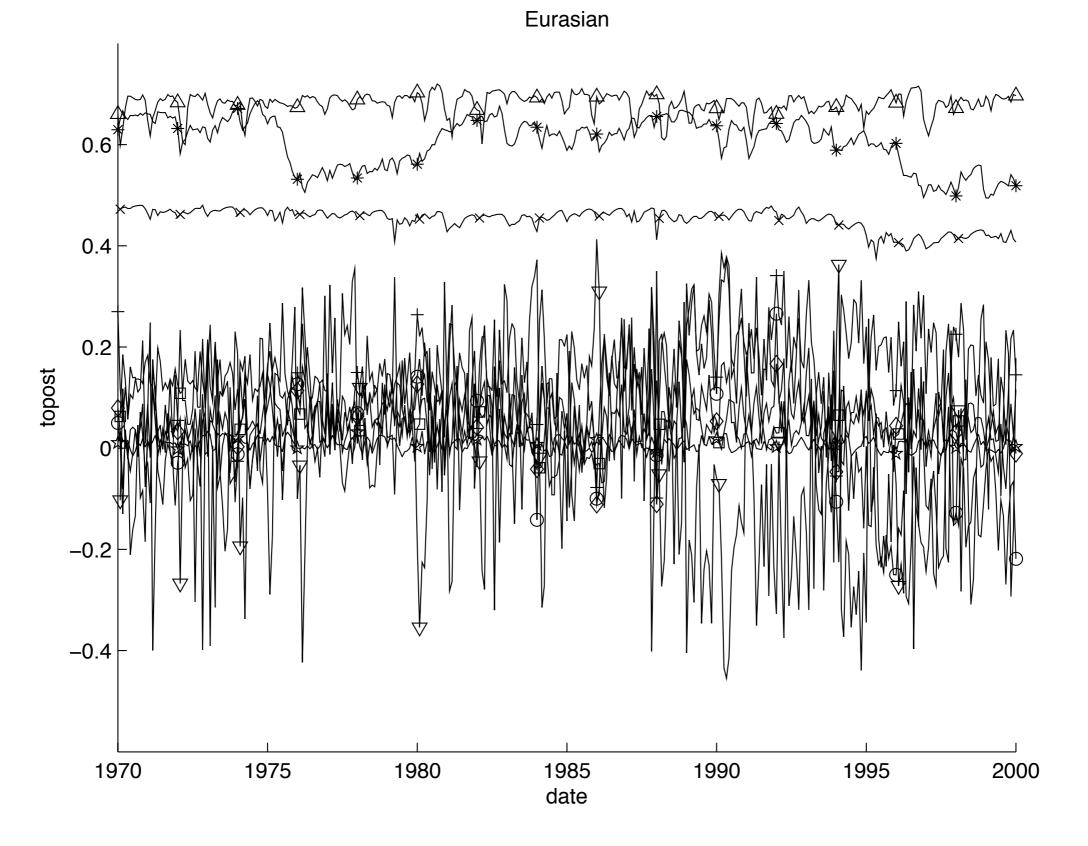


Three illustrations: neptune vertical viscosity sea ice

neptune (level I):
$$\partial \mathbf{Y}_{\partial t} = \mathbf{F}(\mathbf{Y}) + \mathbf{G} + \mathbf{C} \cdot \partial_{\mathbf{Y}} H$$

 $\mathbf{C} \cdot \partial_{\mathbf{Y}} H \approx \mathbf{C} \cdot \partial_{\mathbf{Y}} \partial_{\mathbf{Y}} H \cdot (\mathbf{Y}^* - \mathbf{Y}) \equiv \mathbf{K} \cdot (\mathbf{Y}^* - \mathbf{Y})$
 $\mathbf{u}^* = -L^2 \mathbf{f} \times \nabla \log D$ (QG energy & enstrophy)
Replace $\partial_t \mathbf{u} + ... = \partial_z A_z \partial_z \mathbf{u} + \nabla_h A_h \nabla_h \mathbf{u} + ...$
with $\partial_t \mathbf{u} + ... = \partial_z A_z \partial_z \mathbf{u} + \nabla_h A_h \nabla_h (\mathbf{u} - \mathbf{u}^*) + ...$

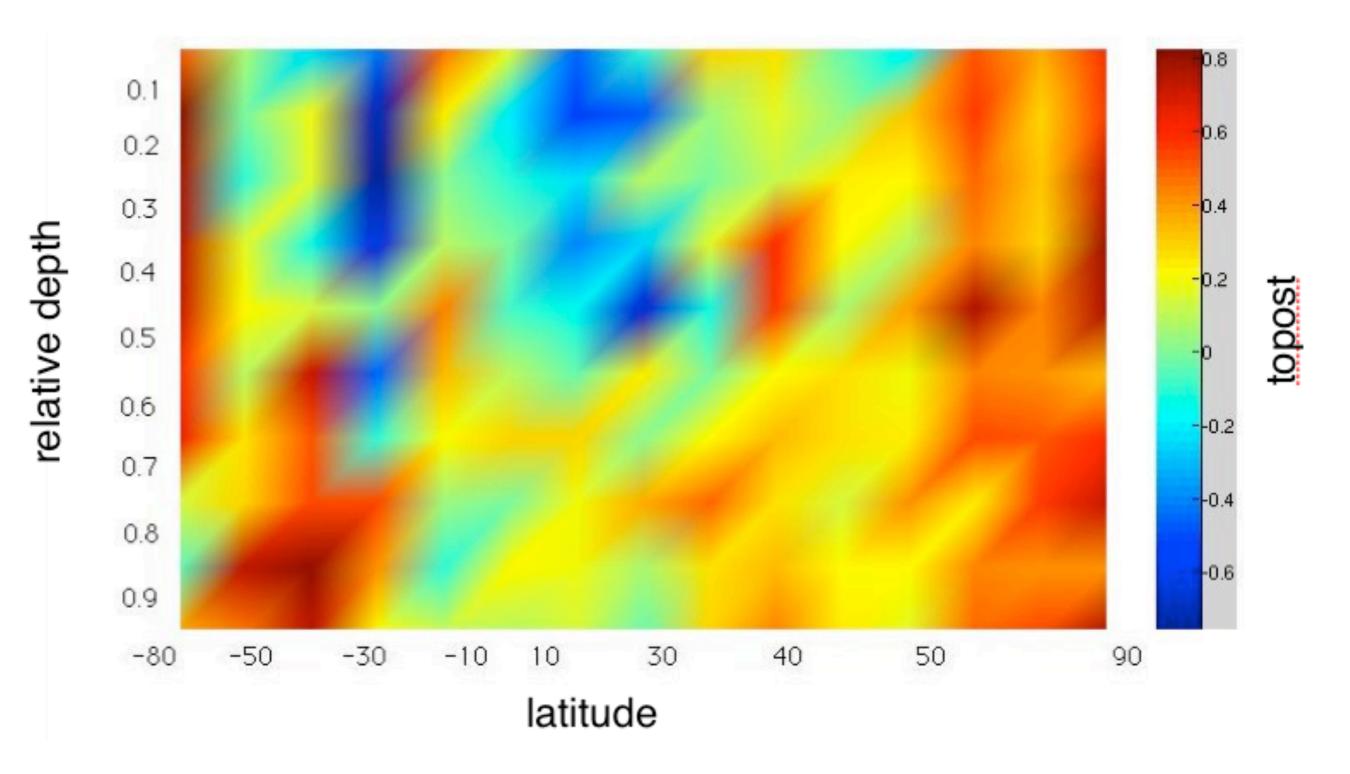




Neptune makes big difference but ... better or worse?

Observed from 17120 current meters over 83087 months (JGR, 2008):

Topostrophy vs. latitude and relative depth



Observations are completely outside the range of traditional modeling????

Three illustrations: neptune vertical viscosity sea ice

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$$\partial \mathbf{Y}_{\partial t} = \mathbf{F}(\mathbf{Y}) + \mathbf{G} + \mathbf{C} \cdot \partial_{\mathbf{Y}} H$$

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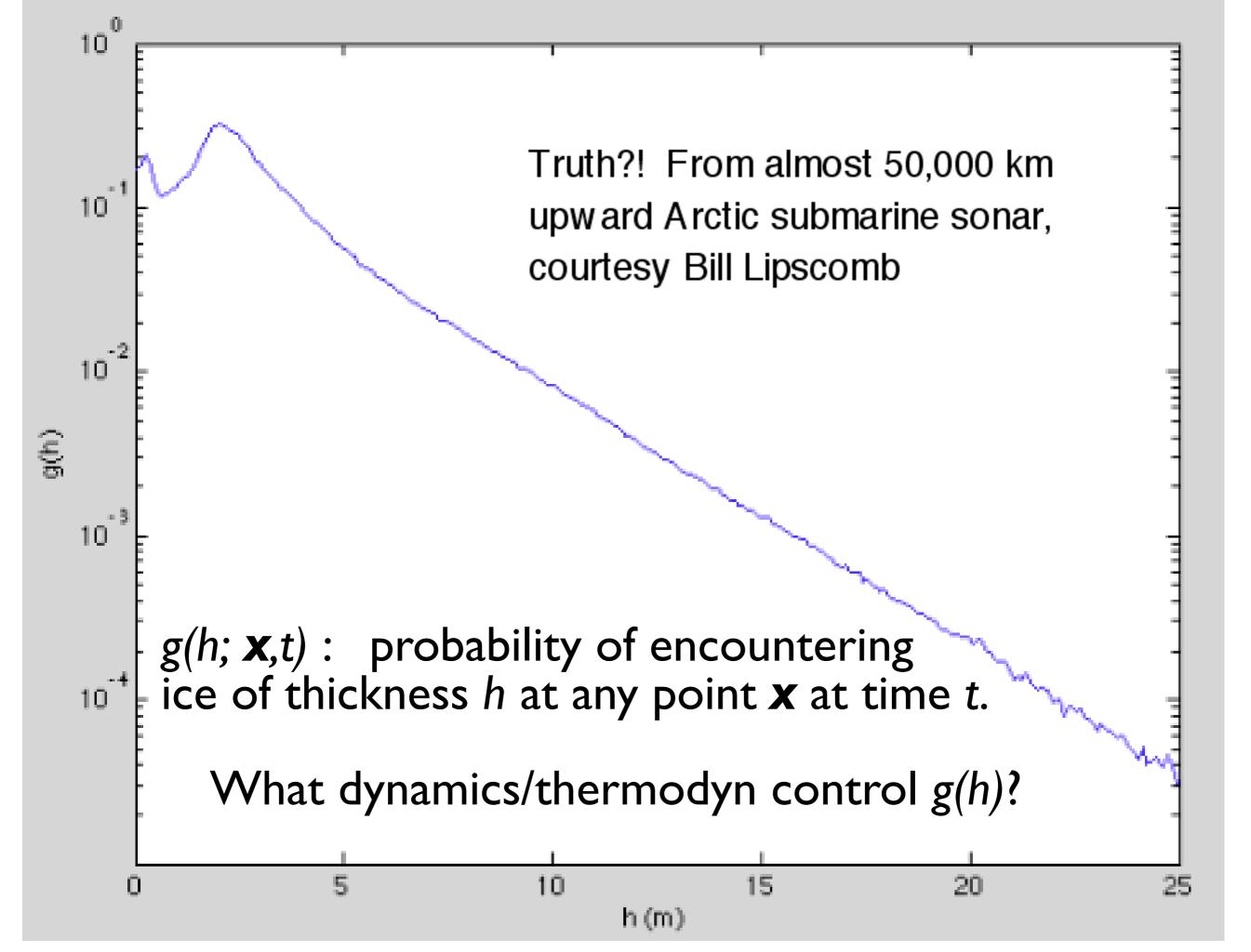
vertical viscosity: $A_z = (f^2/N^2)A_?$

 $f^2/N^2 \approx 10^{-3}$, $A_2 \approx 10^2$, then $A_z \approx 10^{-1}m^2/s$ (not 10^{-4} to 10^{-3})

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Three illustrations:

neptune vertical viscosity sea ice



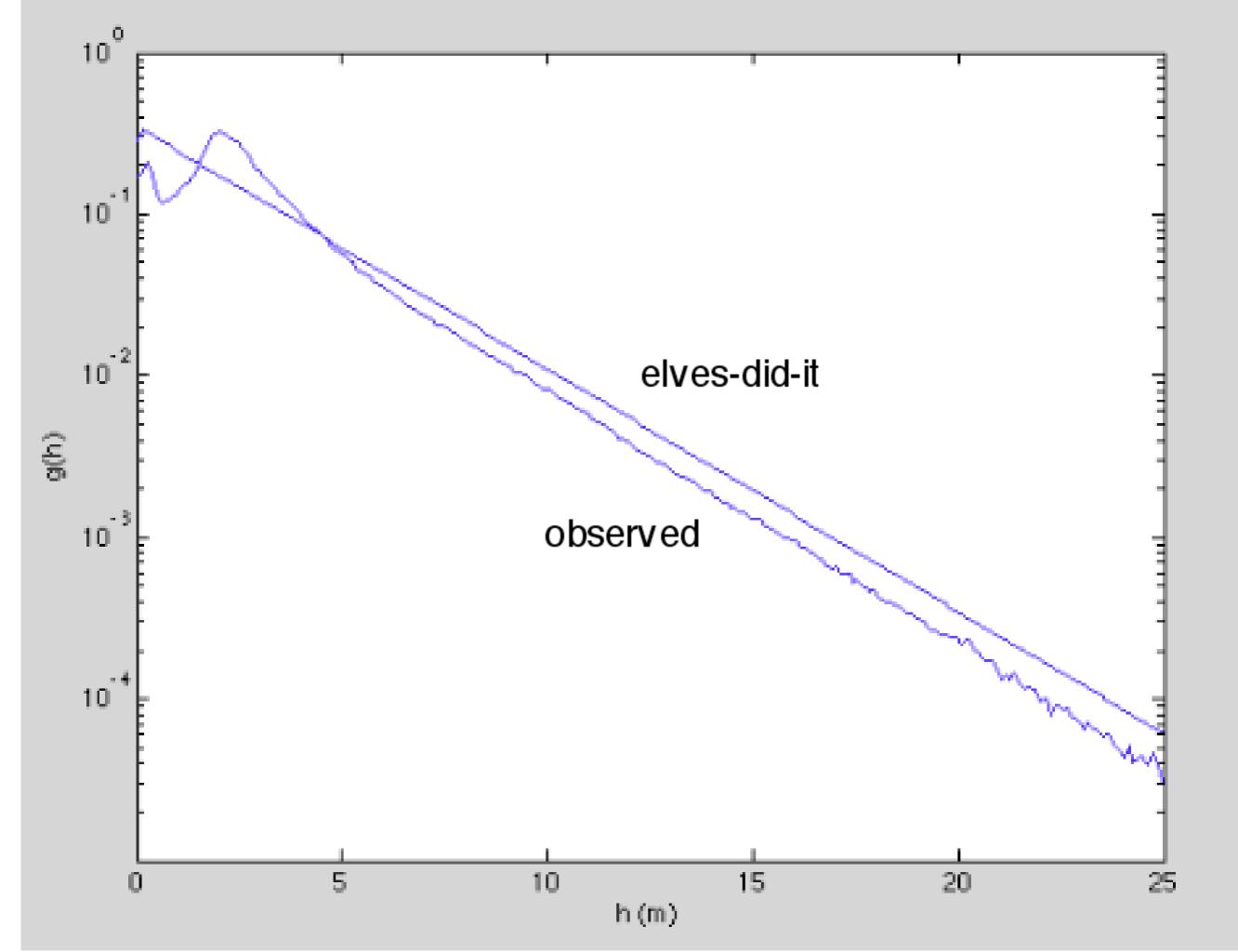


A (very!) little math:

Maximise
$$S = -\int_{0}^{\infty} \log(g)gdh$$

subject to $\int_{0}^{\infty} gdh = 1$ and $\int_{0}^{\infty} ghdh = H$
and *presto!* $g(h) = \exp(-h/H)/H$

Thus, Santa's elves, underoccupied during off-season, randomly toss ice into the Arctic. Is that easy, or what?



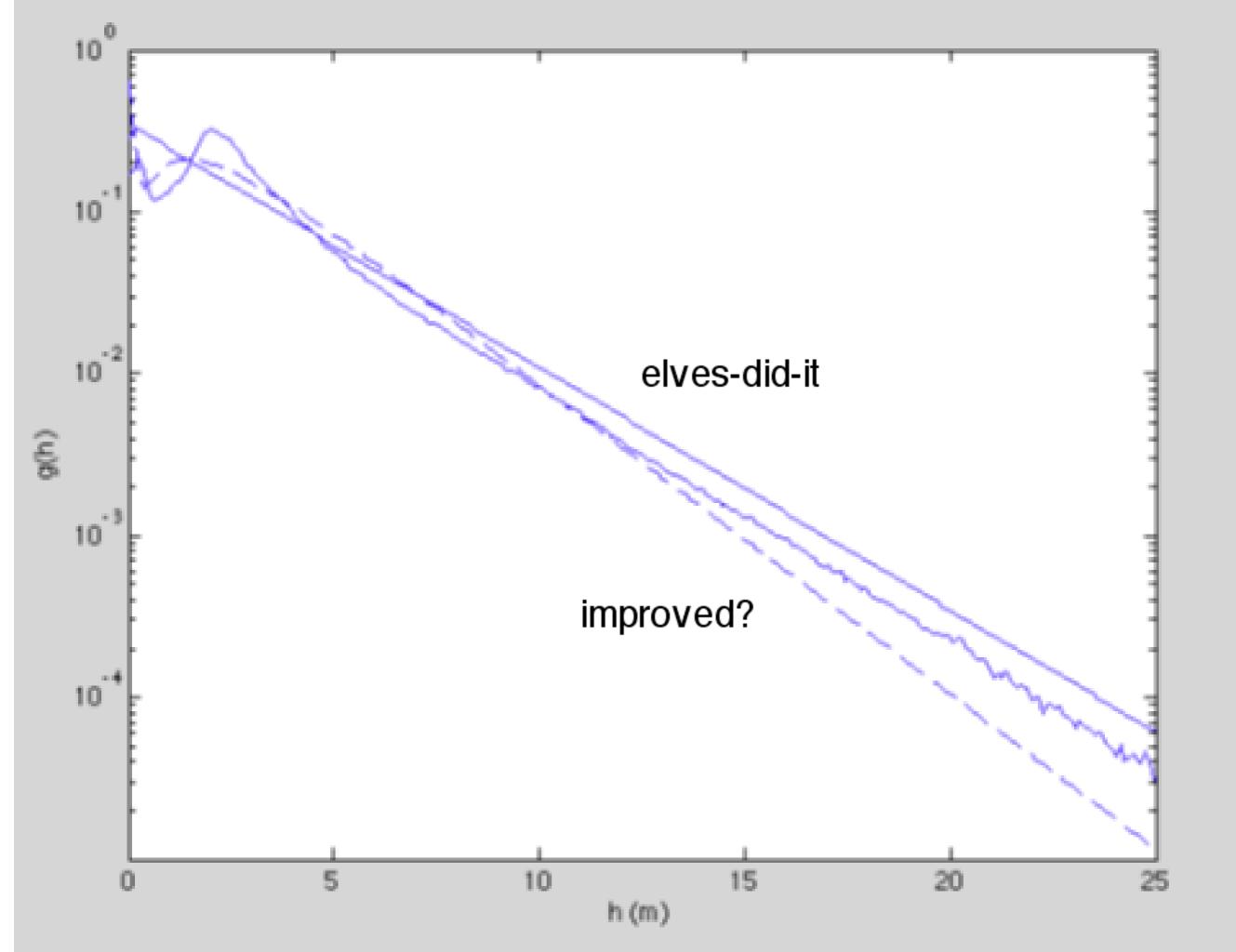
OK, not the greatest ever. But fiddle just a little.

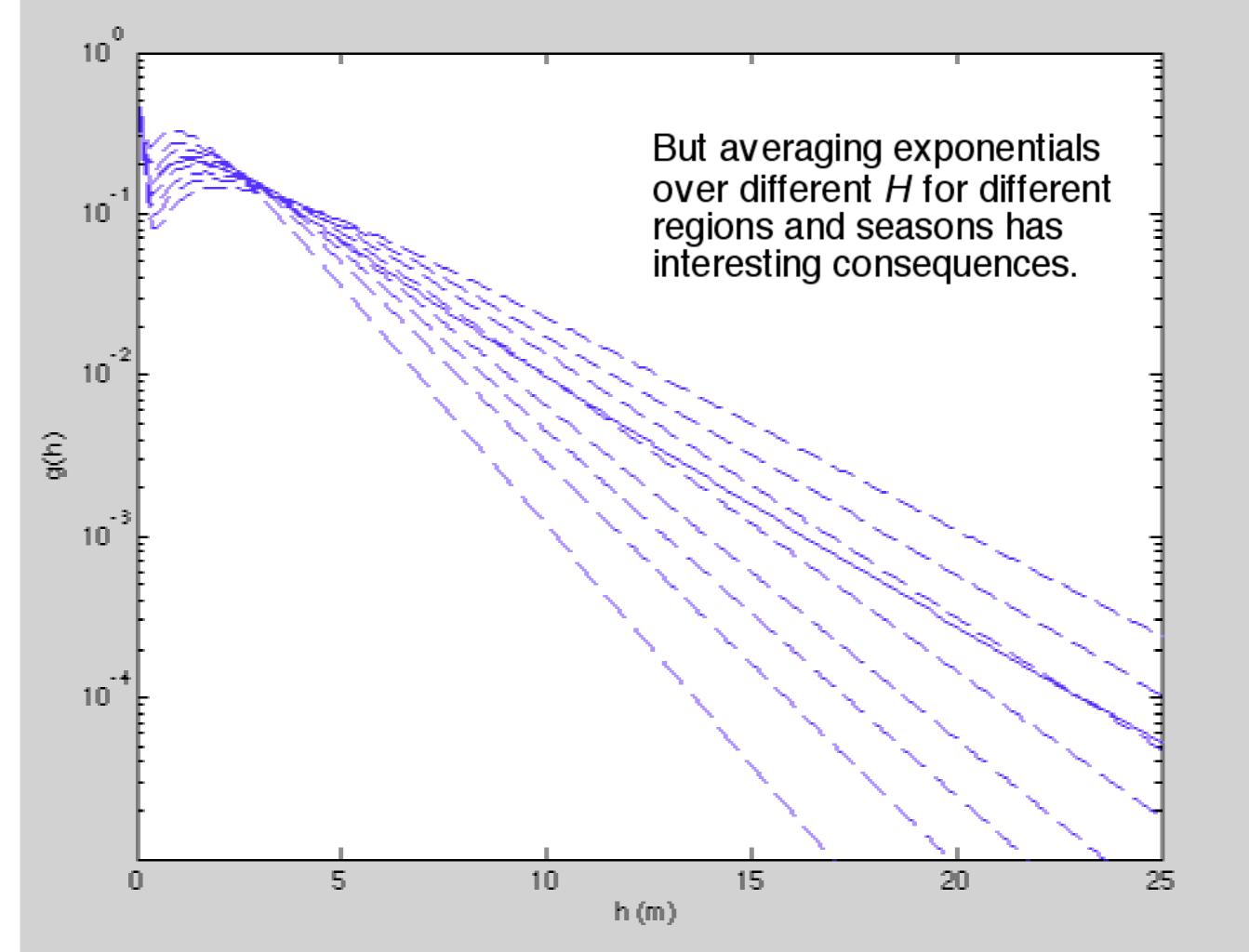
Wind shifts, tides, inertial oscillations, ... open ice.

Recognize an "open water" (thin ice?) fraction 1-A. In leads during freezing, thin new ice rapidly forms but is easily smunched into thicker ice. Fix it?

$$D = \frac{A}{\left(2a-1\right)H} \left(e^{-\frac{h}{aH}} - e^{-\frac{h}{H-aH}}\right) + \frac{1-A}{bH}e^{-\frac{h}{bH}}$$

Improved? Let's see.



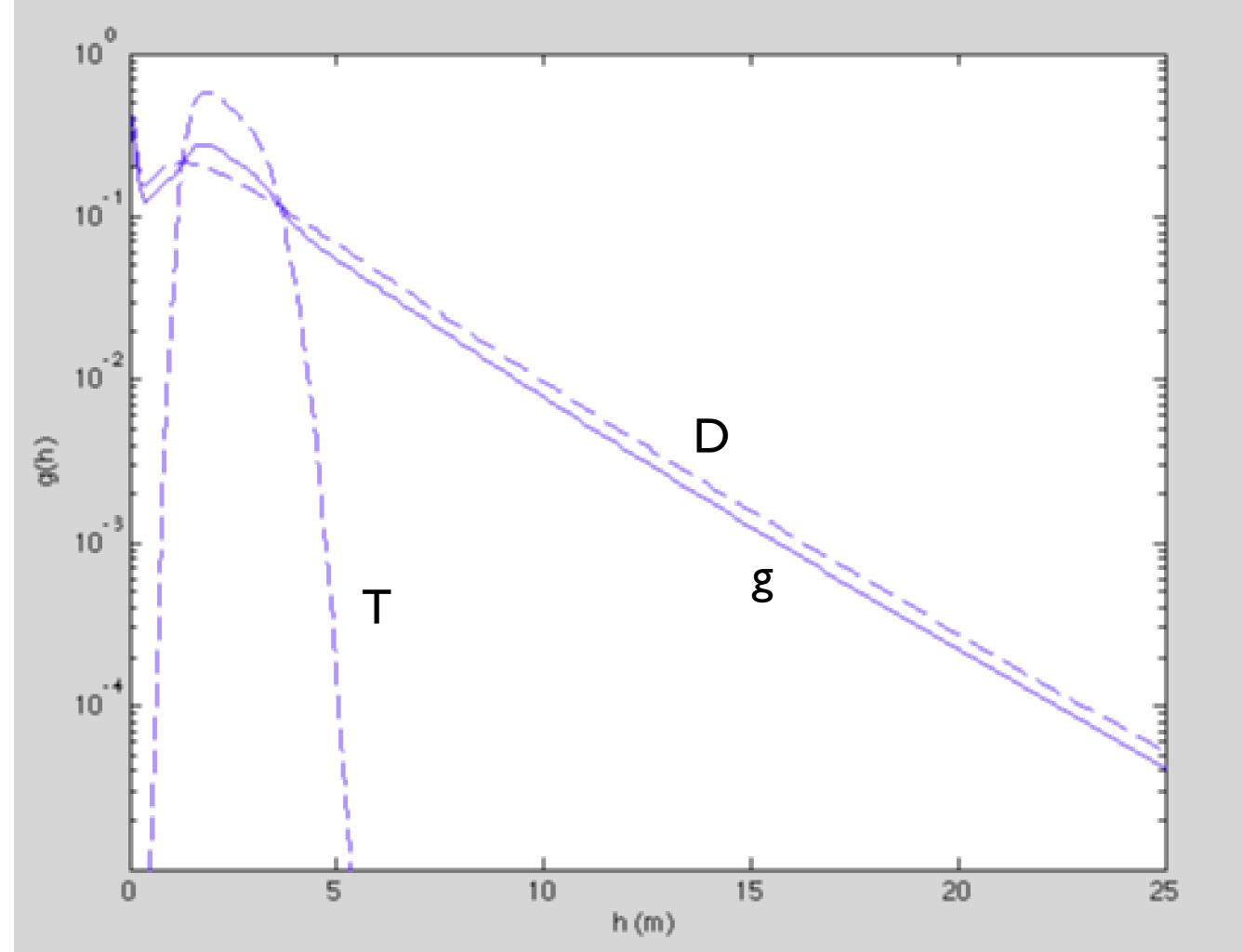


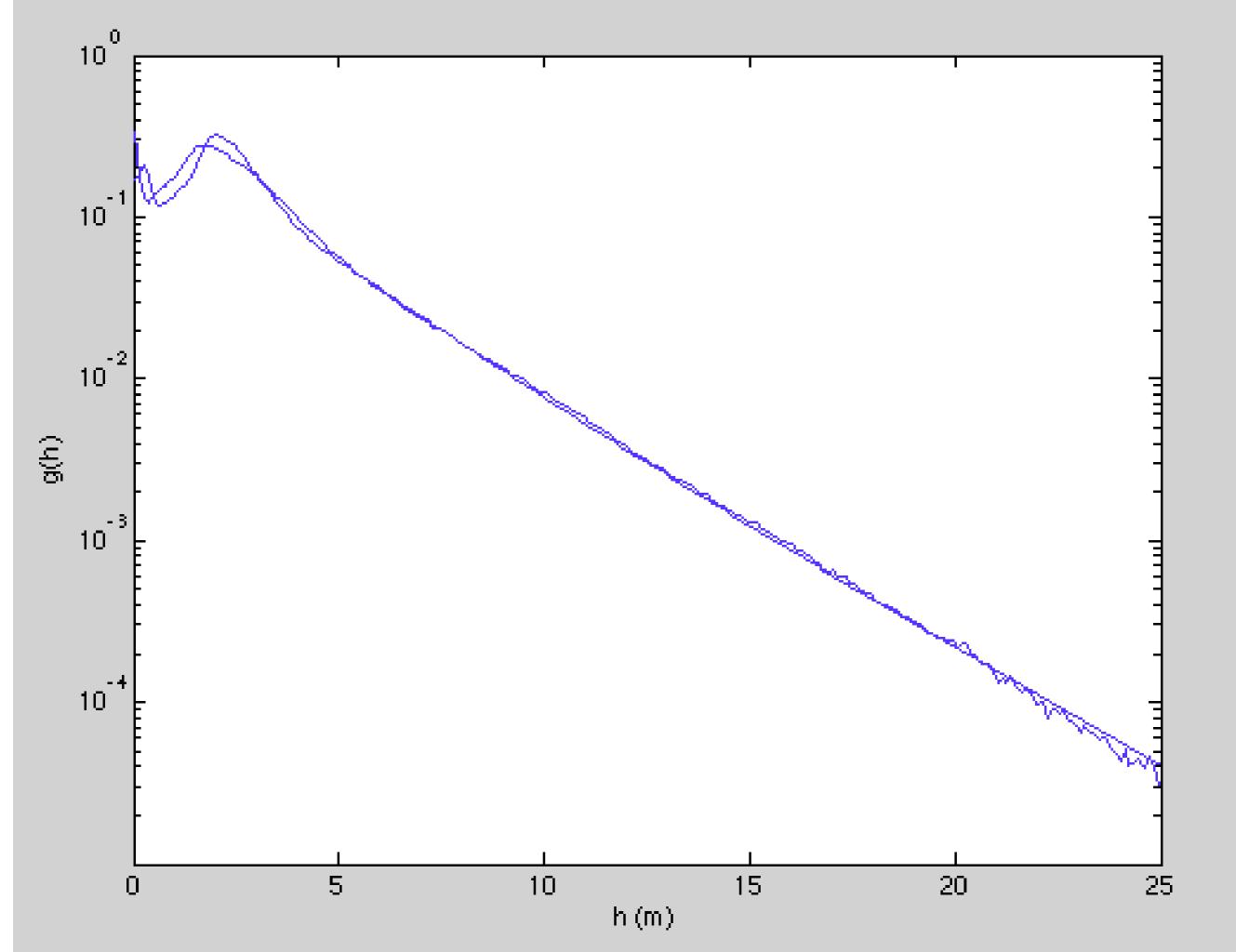
Dynamics -- alone -- spreads out g(h), approaching D.

Thermodynamics -- alone -- would focus g(h), e.g. as

$$T = e^{-\frac{(h-aH)^2}{2b^2}} \sqrt{2\pi b}$$
 which also needs be averaged over

various *H*. Combine to g(h) = cD + (1-c)T for some *c*. See a case *c* =0.8:





This is just fudge factors!

Yes. But not so many. b only need be small, like $b\sim0.1\,\text{m}$. a is constrained, 0.5 < a < 1. I chose a = 0.7. That leaves c. I chose c = 0.8 (more dynamic, not so thermodyn.)

Models solve dA/dt and dH/dt. Fudge something for dc/dt, like integrated deformation? Keep first- and multi-year fractions?

Summary:

Stat mech completes the equations of motion. Includes entropic forcing.

- E.g., neptune
- E.g., vertical viscosity

Entropy also is about economy (intellectual & computational). When outcomes depend upon many detailed interactions, problems get *easier*. Entropy calculus (a crutch) *avoids* difficult stuff. E.g., sea ice thickness